

IPhO 2018 Lisbon, Portugal

Solutions to Theory Problem 2

Where is the neutrino?

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Where is the neutrino? (10 points)

Part A. ATLAS Detector physics (4.0 points)

A.1

The magnetic force is the centripetal force:

$$m\frac{v^2}{r} = evB \Rightarrow r = \frac{mv}{eB}$$

First express the velocity in terms of the kinetic energy,

$$K = \frac{1}{2}mv^2 \Rightarrow v = \sqrt{\frac{2K}{m}},$$

and then insert it in the expression above for the radius to get

A.1

 $r = \frac{\sqrt{2Km}}{eB} \,.$

0.5pt

A.2

The radius of the circular motion of a charged particle in the presence of a uniform magnetic field is given by,

$$r = \frac{mv}{eB}$$
.

This formula is valid in the relativistic scenario if the mass correction, $m \to \gamma m$ is included:

$$r = \frac{\gamma m v}{eB} = \frac{p}{eB} \Rightarrow p = reB \,.$$

Note that the radius of the circular motion is half the radius of the inner part of the detector. One obtains [1 MeV/ $c=5.34\times10^{-22}$ m kg s $^{-1}$]

A.2

0.5pt

$$p = 330 \text{ MeV/}c$$
.

A.3

The acceleration for the particle is $a=rac{evB}{\gamma m}\simrac{ecB}{\gamma m}$, in the ultrarelativistic limit. Then,

$$P = \frac{e^4 c^2 \gamma^4 B^2}{6\pi \epsilon_0 c^3 \gamma^2 m^2} = \frac{e^4 \gamma^2 c^4 B^2}{6\pi \epsilon_0 c^5 m^2}.$$

Since $E=\gamma mc^2$ we can obtain $\gamma^2c^4=\frac{E^2}{m^2}$ and, finally,

$$P = \frac{e^4}{6\pi\epsilon_0 m^4 c^5} E^2 B^2.$$

Therefore,

A.3

$$\xi=\frac{1}{6\pi}\,,\quad n=5\ \ \text{and}\quad k=4.$$

1.0pt

A.4

The power emitted by the particle is given by,

$$P = -\frac{{\rm d}E}{{\rm d}t} = \frac{e^4}{6\pi\epsilon_0 m^4 c^5} E^2 B^2 \,.$$

The energy of the particle as a function of time can be calculated from

$$\int_{E_0}^{E(t)} \frac{1}{E^2} \mathrm{d}E = - \int_0^t \frac{e^4}{6\pi \epsilon_0 m^4 c^5} B^2 \mathrm{d}t \,,$$

where $E(0) = E_0$. This leads to,

$$\frac{1}{E(t)} - \frac{1}{E_0} = \frac{e^4 B^2}{6\pi\epsilon_0 m^4 c^5} \; t \quad \Rightarrow \quad E(t) = \frac{E_0}{1 + \alpha E_0 t} \, , \label{eq:energy}$$

with

A.4

$$\alpha = \frac{e^4 B^2}{6\pi \epsilon_0 m^4 c^5} \,. \label{eq:alpha}$$

1.0pt

A.5

If the initial energy of the electron is 100 GeV, the radius of curvature is extremely large ($r=\frac{E}{eBc}\approx 167$ m). Therefore, in approximation, one can consider the electron is moving in the inner part of the ATLAS detector along a straight line. The time of flight of the electron is t=R/c, where R=1.1 m is the radius of the inner part of the detector. The total energy lost due to synchrotron radiation is,

$$\Delta E = E(R/c) - E_0 = \frac{E_0}{1 + \alpha E_0 \frac{R}{c}} - E_0 \approx -\alpha E_0^2 \frac{R}{c}$$

and

A.5

0.5pt

$$\Delta E = -56~\mathrm{MeV}$$
 .

A.6

In the ultrarelativistic limit, $v \approx c$ and $E \approx pc$. The cyclotron frequency is,

$$\omega(t) = \frac{c}{r(t)} = \frac{ecB}{p(t)} = \frac{ec^2B}{E(t)}$$

$$\omega(t) = \frac{ec^2B}{E_0} \left(1 + \frac{e^4B^2}{6\pi\epsilon_0 m^4c^5} E_0 t \right) \,.$$

0.5pt

B.1

Since the W^+ boson decays into an anti-muon and a neutrino, one can use principles of conservation of energy and linear momentum to calculate the unknown $p_{\rm z}^{(\nu)}$ of the neutrino. Moreover, the anti-muon and the neutrino can be considered massless, which implies that the magnitude of their momenta (times c) and their energies are the same. Therefore, the conservation of linear momentum can be expressed as

$$\vec{p}^{(W)} = \vec{p}^{(\mu)} + \vec{p}^{(\nu)},$$

and the conservation of energy as,

$$E^{(W)} = cp^{(\mu)} + cp^{(\nu)}$$
.

In addition, one can also relate the energy and the momentum of the \mathcal{W}^+ boson through its mass,

$$m_W^2 = (E^{(W)})^2/c^4 - (p^{(W)})^2/c^2$$

which leads to a quadratic equation on $p_z^{(\nu)}$,

Part B. Finding the neutrino (6.0 points)

$$\begin{array}{lll} m_W^2 & = & \left[(p^{(\mu)} + p^{(\nu)})^2 - (\vec{p}^{\;(\mu)} + \vec{p}^{\;(\nu)})^2 \right]/c^2 \\ & = & \left(2p^{(\mu)}p^{(\nu)} - 2\vec{p}^{\;(\mu)} \cdot \vec{p}^{\;(\nu)} \right)/c^2 \end{array}$$

B.1

$$m_W^2 = \frac{1}{c^2} \left(2 p^{(\mu)} \sqrt{(p_{\rm T}^{(\nu)})^2 + (p_z^{(\nu)})^2} - 2 \vec{p}_{\rm T}^{\;(\mu)} \cdot \vec{p}_{\rm T}^{\;(\nu)} - 2 p_z^{(\mu)} p_z^{(\nu)} \right) \,. \label{eq:mw}$$

1.5pt

B.2

The numerical substitution directly in the answer of B.1, using

$$p^{(\mu)} = 37.2\,\mathrm{GeV}/c \hspace{0.5cm} m_W^2 c^2 = 6464.2\,(\mathrm{GeV}/c)^2 \hspace{0.5cm} p_\mathrm{T}^{(\nu)~2} = 10\,864.9\,(\mathrm{GeV}/c)^2$$

$$\vec{p}_{\rm T}^{\;(\mu)} \cdot \vec{p}_{\rm T}^{\;(\nu)} = 2439.3 ({\rm GeV}/c)^2 \qquad p_z^{(\mu)} = -12.4 \, {\rm GeV}/c \,, \label{eq:pt_T}$$

leads to

$$6464.2 = 74.4 \sqrt{10\,864.9 + p_z^{(\nu)\,\,2}} - 4878.6 + 24.8 p^{(\nu)}\,.$$

This is a quadratic equation, equivalent to

$$0.88889 \, p_z^{(\nu) \, 2} + 101.64 \, p_z^{(\nu)} - 12378 = 0$$

whose solutions are:

B.2

1.5pt

$$p_z^{(\nu)} = 74.0 \text{ GeV/}c$$

$$p_z^{(\nu)} = 74.0~{\rm GeV}/c \qquad {\rm or} \qquad p_z^{(\nu)} = -188.3~{\rm GeV}/c. \label{eq:pz}$$

The general solution of the equation above in B.1 leads to

$$\begin{array}{lcl} p_z^{(\nu)} & = & \frac{2\vec{p}_{\mathsf{T}}^{\;(\mu)} \cdot \vec{p}_{\mathsf{T}}^{\;(\nu)} p_z^{(\mu)} + m_W^2 c^2 p_z^{(\mu)}}{2(p_{\mathsf{T}}^{(\mu)})^2} \\ \\ & \pm \frac{p^{(\mu)} \sqrt{-4(p_{\mathsf{T}}^{\;(\mu)})^2 (p_{\mathsf{T}}^{\;(\nu)})^2 + 4(\vec{p}_{\mathsf{T}}^{\;(\mu)} \cdot \vec{p}_{\mathsf{T}}^{\;(\nu)})^2 + 4\vec{p}_{\mathsf{T}}^{\;(\mu)} \cdot \vec{p}_{\mathsf{T}}^{\;(\nu)} m_W^2 c^2 + m_W^4 c^4}{2(p_{\mathsf{T}}^{\;(\mu)})^2} \end{array}$$

Numerical substitution leads to the above mentioned values for $p_z^{(\nu)}$.

B.3

The final state particles of the top quark decay are the anti-muon, the neutrino and jet 1. Since the neutrino is now fully reconstructed the energy and linear momentum of the top quark can be calculated

$$\begin{array}{lcl} E^{({\rm t})} & = & cp^{(\mu)} + cp^{(\nu)} + cp^{(j_1)} \\ \vec{p}^{~({\rm t})} & = & \vec{p}^{~(\mu)} + \vec{p}^{~(\nu)} + \vec{p}^{~(j_1)} \,. \end{array}$$

The top quark mass is,

$$\begin{array}{lcl} m_{\mathsf{t}} & = & \sqrt{(E^{(\mathsf{t})})^2/c^4 - (\vec{p}^{\;(\mathsf{t})})^2/c^2} \\ \\ & = & c^{-1}\sqrt{\left(p^{(\mu)} + p^{(\nu)} + p^{(j_1)}\right)^2 - \left(\vec{p}^{\;(\mu)} + \vec{p}^{\;(\nu)} + \vec{p}^{\;(j_1)}\right)^2} \,. \end{array}$$

The substitution of values leads to two possible masses:

B.3

$$m_{\rm t} = 169.3~{\rm GeV}/c^2$$

$$m_{\rm t} = 169.3~{\rm GeV}/c^2 \qquad {\rm or} \qquad m_{\rm t} = 311.2~{\rm GeV}/c^2 \label{eq:mt}$$

1.0pt

B.4

According to the frequency distribution for signal (dashed line), the probability of the $m_{\rm t}=169.3~{\rm GeV}/c^2$ solution is roughly 0.1 while the probability of the $m_{\rm t}=311.2~{\rm GeV/c^2}$ solution is below 0.01. Therefore,

The most likely candidate is the $m_{\rm t}=169.3~{\rm GeV/c^2}$ solution. **B.4**

1.0pt

B.5

The top quark energy for the most likely candidate is $E^{(t)}=cp^{(\mu)}+cp^{(\nu)}+cp^{(j_1)}=272.6~{
m GeV}$.

$$d = vt = v\gamma t_0 = \frac{p^{(\mathsf{t})}}{m_{\mathsf{t}}} t_0 = ct_0 \sqrt{\frac{{E^{(\mathsf{t})}}^2}{m_{\mathsf{t}}^2 c^4} - 1} \,.$$

B.5
$$d = 2 \times 10^{-16} \ \mathrm{m} \, .$$